# A stable scheme and its convergence analysis for a $2 D$ dynamic Q-tensor model of nematic liquid crystals 

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#### Abstract

We propose an unconditionally stable numerical scheme for a $2 D$ dynamic $Q$-tensor model of nematic liquid crystals. This dynamic $Q$-tensor model is a $L^{2}$ gradient flow generated by the liquid crystal free energy that contains a cubic term, which is physically relevant but makes the free energy unbounded from below, and for this reason, has been avoided in other numerical studies. The unboundedness of the energy brings significant difficulty in analyzing the model and designing numerical schemes. By using a stabilizing technique, we construct an unconditionally stable scheme, and establish its unique solvability and convergence. Our convergence analysis also leads to, as a byproduct, the well-posedness of the original PDE system for the 2D Q-tensor model. Several numerical examples are presented to validate and demonstrate the effectiveness of the scheme.


## 1 Introduction

Liquid crystals are an intermediate state of matter between the commonly observed solid and liquid that has no or partial positional order but do exibit an orientional order. The nematic phase is the simplest among all liquid crystal phases whose rod-like molecules have no translational order but possesses a certain degree of long-range orientaional ordering. The Landau-de Gennes theory [7] is a continuum theory to describe the nematic liquid crystals. In this framework, it is widely accepted that the local orientation and degree of ordering for the liquid crystal molecules are characterized by a symmetric, traceless $d \times d$ tensor called the $Q$-tensor in $\mathbb{R}^{d}(d=1,2,3)[1,22]$. The $Q$-tensor vanishes in the isotropic phase, and hence it serves as an order parameter. The $Q$-tensor order parameter may exhibit two different phases, namely the uniaxial phase and the biaxial phase. In the former phase, $Q$ has uniaxial symmetry and the symmetry axis is defined by a unit vector $\vec{n}$ called the director. In the latter biaxial phase, the structure of $Q$ is more complicated. There exists a vast literature on the mathematical study of the Landau-de Gennes theory, see [2,3,9,18,21,24,25] and the references therein.

The equilibrium states are physically observable configurations which correspond to either global or local minimizers of the free energy subject to certain imposed boundary conditions. Consider a

[^0]nematic liquid crystal filling a smooth, bounded domain $\Omega \subset \mathbb{R}^{d}$, and for the sake of simplicity, we suppose that the material is spatially homogeneous and the temperature is constant. Historically, the first step toward the understanding of its free energy is attributed to [11,23] where the free energy density functional (called the Oseen-Frank energy density) is expressed in terms of the director $\vec{n}$ with elastic constants $K_{1}, \cdots K_{4}$ :
\[

$$
\begin{equation*}
W_{O F}=\frac{K_{1}}{2}(\nabla \cdot \vec{n})^{2}+\frac{K_{2}}{2}|\vec{n} \times(\nabla \times \vec{n})|^{2}+\frac{K_{3}}{2}|\vec{n} \cdot(\nabla \times \vec{n})|^{2}+\frac{K_{2}+K_{4}}{2}\left[\operatorname{tr}(\nabla \vec{n})^{2}-(\nabla \cdot \vec{n})^{2}\right] . \tag{1.1}
\end{equation*}
$$

\]

Here $K_{1}, \cdots K_{3}$ measure [11] the resistance of three basic distortions, called splay, twist and bend, respectively, and the last term in (1.1) is related to the twisted splay distortion, which is a null Lagrange term but is kept in most literature because this term does contribute to the total free energy for some types of boundary value problems [14]. The Oseen-Frank formulation is generally consistent with experiment except near the nematic-smectic phase transition [8]. In order to generalize the Oseen-Frank description close to the clearing point, de Gennes [7] proposed a Ginzburg-Landau type expansion of the free energy in terms of the tensor parameter $Q$ and its spatial derivatives. The Landau-de Gennes free energy functional is derived as a nonlinear integral functional of the $Q$-tensor and its spatial derivatives [1]:

$$
\begin{equation*}
\mathcal{E}[Q]=\int_{\Omega} \mathcal{F}(Q(x)) d x \tag{1.2}
\end{equation*}
$$

where $Q$ is in the $Q$-tensor space (c.f. [1,3,22]) defined by

$$
\mathcal{S}^{(d)} \stackrel{\text { def }}{=}\left\{M \in \mathbb{R}^{d \times d} \mid \sum_{i=1}^{d} M^{i i}=0, M^{i j}=M^{j i} \in \mathbb{R}, \forall i, j=1, \cdots, d\right\} .
$$

The free energy density functional $\mathcal{F}$ consists of the elastic part $\mathcal{F}_{e l}$ that depends on the gradient of $Q$, and the bulk part $\mathcal{F}_{\text {bulk }}$ that depends on $Q$ only [15], i.e.,

$$
\begin{equation*}
\mathcal{F}(Q)=\mathcal{F}_{\text {el }}+\mathcal{F}_{\text {bulk }} . \tag{1.3}
\end{equation*}
$$

The bulk free-energy density $\mathcal{F}_{\text {bulk }}$ is typically a truncated expansion in the scalar invariants of the tensor $Q$. In the simplest setting one may take

$$
\begin{equation*}
\mathcal{F}_{\text {bulk }} \xlongequal{\text { def }} \frac{a}{2} \operatorname{tr}\left(Q^{2}\right)+\frac{b}{3} \operatorname{tr}\left(Q^{3}\right)+\frac{c}{4} \operatorname{tr}^{2}\left(Q^{2}\right), \tag{1.4}
\end{equation*}
$$

where $a, b, c$ are assumed to be bulk material constants. This bulk term (1.4) embodies the ordering/disordering effects, which drive the nematic-isotropic phase transition. It depends only on the eigenvalues of $Q$. Meaningful simulations can be performed using an expansion truncated at the fourth order, to which we have to use in order to have a potential with multiple stable local minima [9].

On the other hand, the elastic free-energy density $\mathcal{F}_{\text {el }}$ gives the strain energy density due to spatial variations in the tensor order parameter. Its simplest form that is invariant under rigid rotations and material symmetry is as follows [1,22]:

$$
\begin{equation*}
\mathcal{F}_{e l} \stackrel{\text { def }}{=} L_{1}|\nabla Q|^{2}+L_{2} \partial_{j} Q^{i k} \partial_{k} Q^{i j}+L_{3} \partial_{j} Q^{i j} \partial_{k} Q^{i k}+L_{4} Q^{l k} \partial_{k} Q^{i j} \partial_{l} Q^{i j} . \tag{1.5}
\end{equation*}
$$

Here and after we use the Einstein summation convention over repeated indices. The material elastic constants $L_{k}(k=1,2,3,4)$ are assumed to be non-dimensional. It is worth pointing out that $\mathcal{F}_{e l}$ in (1.5) consists of three independent terms with constants $L_{1}, L_{2}, L_{3}$ that are quadratic in the first partial derivatives of the components of $Q$, plus an unusual cubic term with constant $L_{4}$. As mentioned in [15,18], the retention of this $L_{4}$ cubic term is due to the consideration that it gives a complete reduction of $\mathcal{F}[Q]$ to the classical Oseen-Frank energy density $W_{O F}$. This is done by formally taking $Q(x)=s_{+}\left(\vec{n}(x) \otimes \vec{n}(x)-\frac{1}{d} \mathbb{I}\right)$, where $s_{+} \in \mathbb{R}^{+}$and substituting it in (1.1). It is shown in $[15,18]$ that if $L_{4}=0$, then $K_{1} \equiv K_{3}$ during the reduction, which clearly contradicts experiment. On the other hand, this $L_{4}$ term causes the Landau-de Gennes free energy to be unbounded from below [2].

In order to remedy the aforementioned deficiency in the static configurations, one way is to replace the bulk potential part defined in (1.5) with a singular type potential [2]; alternatively, a dynamic case is later proposed in [15] to keep the more common bulk potential in (1.5). More specifically, the authors in [15] studies the following $L^{2}$ gradient flow in $\mathbb{R}^{2}$ corresponding to the energy functional $\mathcal{E}[Q]$ where $Q$ takes values in $\mathcal{S}^{(2)}$ :

$$
\begin{equation*}
\frac{\partial Q^{i j}}{\partial t}=-\left(\frac{\delta \mathcal{E}}{\delta Q}\right)^{i j}+\lambda \delta^{i j}+\mu^{i j}-\mu^{j i}, \quad 1 \leq i, j \leq 2 . \tag{1.6}
\end{equation*}
$$

In (1.6), $\lambda$ is the Lagrange multiplier corresponding to the traceless constraint and $\mu=\left(\mu^{i j}\right)_{2 \times 2}$ is the Lagrange multiplier corresponding to the matrix symmetry constraint, and $\frac{\delta \mathcal{E}}{\delta Q}$ denotes the variational derivative of $\mathcal{E}$ with respect to $Q$. In addition, we always impose hereafter the coercivity condition [15] (see also [9] for its counterpart in $3 D$ )

$$
\begin{equation*}
L_{1}+L_{2}>0, \quad L_{1}+L_{3}>0, \tag{1.7}
\end{equation*}
$$

and

$$
\begin{equation*}
c>0 . \tag{1.8}
\end{equation*}
$$

From the modeling point of view, (1.7) is imposed to guarantee that the summation of the first three quadratic terms concerning $L_{1}, L_{2}, L_{3}$ in $\mathcal{F}_{e l}$ is positive definite, while (1.8) is to ensure $\mathcal{F}_{\text {bulk }}$ is bounded from below. Moveover, as noted in [5,15], the term $\frac{b}{3} \operatorname{tr}\left(Q^{3}\right)$ can be ignored from (1.4) since $\operatorname{tr}\left(Q^{3}\right)=0$ for any $Q \in \mathcal{S}^{(2)}$.

After expansion, the evolution equation (1.6) reads

$$
\begin{equation*}
\partial_{t} Q^{i j}=\zeta \Delta Q^{i j}+L_{4}\left\{2 \partial_{k}\left(Q^{l k} \partial_{l} Q^{i j}\right)-\partial_{i} Q^{k l} \partial_{j} Q^{k l}+\frac{|\nabla Q|^{2} \delta^{i j}}{2}\right\}-\left[a+c \operatorname{tr}\left(Q^{2}\right)\right] Q^{i j} \tag{1.9}
\end{equation*}
$$

for $1 \leq i, j \leq 2$, with initial and boundary conditions given by

$$
\begin{equation*}
Q(x, 0)=Q^{0}(x), \quad \text { and }\left.\quad Q(x, t)\right|_{\partial \Omega}=\tilde{Q}(x),\left.\quad Q^{0}\right|_{\partial \Omega}=\tilde{Q} \tag{1.10}
\end{equation*}
$$

Note that

$$
\begin{equation*}
\zeta \stackrel{\text { def }}{=} 2 L_{1}+L_{2}+L_{3}>0 \tag{1.11}
\end{equation*}
$$

under the coercivity condition (1.7).

Since the free energy $\mathcal{E}[Q]$ is unbounded from below when $L_{4} \neq 0$, generally one may not expect a global existence result to the problem (1.9)-(1.10) without involving a smallness assumption of $Q(\cdot, t)$. To be more precise, this gradient flow gives us the following energy dissipative law for smooth solutions $Q(\cdot, t)$ that satisfies

$$
\frac{d}{d t} \mathcal{E}[Q]=-\int_{\Omega}\left|\frac{\delta \mathcal{E}}{\delta Q}-\lambda \mathbf{I}_{2}+\mu-\mu^{T}\right|^{2} d x
$$

which immediately produces the integral equality

$$
\begin{equation*}
\mathcal{E}[Q(\cdot, t)]+\int_{0}^{t} \int_{\Omega}\left|\frac{\delta \mathcal{E}}{\delta Q}-\lambda \mathbf{I}_{2}+\mu-\mu^{T}\right|^{2} d x d s=\mathcal{E}[Q(\cdot, 0)], \quad \forall t>0 . \tag{1.12}
\end{equation*}
$$

Here $\mathbf{I}_{2}$ stands for the $2 \times 2$ identity matrix. However, we cannot get any a priori control of $\|Q(\cdot, t)\|_{H^{1}(\Omega)}$ from (1.12) because of the unboundedness of $\mathcal{E}[Q]$.

Fortunately, the mathematical structure of (1.9) is exploited thoroughly in [15] so that for any smooth solutions to the evolution problem, the smallness of $\left\|Q_{0}\right\|_{L^{\infty}(\Omega)}$ will be preserved as time evolves. Based on this property plus the coercivity condition (1.7), the authors in [15] obtain the necessary a priori bounds from the energy equality (1.12), which paves the way to obtain the global existence result.

Along the numerical front, there exists only a few studies on the the Q-tensor model. For the stationary cases with $L_{4}=0$, there have been several studies on phase transitions [17], density variations [27], singularities [4] and liquid crystal alignments [6]. For the dynamic Q-tensor model with $L_{4}=0$, a spectral method was used in [30] to study the instability of nanorod dispersions, an adaptive moving mesh method was proposed in [16], and a stable finite element discretization was introduced in [4] for the gradient flow dynamics with constant orientational order parameter. However, to the best of our knowledge, there has been no study, simulation or numerical analysis for the Q-tensor model in the general case with $L_{4} \neq 0$. Notice that this unusual cubic term ( $L_{4} \neq 0$ ) corresponds to the compatibility between the Q-tensor model and the Oseen-Frank model for liquid crystal $[15,18]$.

In this paper we construct an unconditionally stable numerical scheme for the full dynamic Q-tensor model (1.9)-(1.10). Since the system admits an energy law (1.12), it is desirable to design an energetically stable scheme to approximate the Q-tensor model (1.9) -(1.10). On the other hand, the energy stability (or the energy boundedness) does not imply well-posedness of the evolution problem because the non-zero term $L_{4} \neq 0$, unless the $L^{\infty}$ norm of the solution is kept small. Inspired by this observation, we need to show, in addition to energy stability, that $L^{\infty}$ norm of the solutions can be kept small, in order to prove the well-posedness of the nonlinear system at each time step. This is much more challenging than establishing the energy stability.

The rest of the paper is organized as follows. In section 2, we present our semi-discrete numerical scheme for (1.9)-(1.10) and establish its unique solvability and convergence. As a byproduct, we obtain the well-posedness of (1.9)-(1.10). We show some numerical tests in section 3, and demonstrate the accuracy and efficiency of our proposed scheme. Finally, some conclusions are drawn in section 4.

We provide below some notations and definitions to be used in the rest of the paper.
For matrices $A, B \in \mathbb{R}^{2 \times 2}$, we define the Frobenius product between $A$ and $B$ by

$$
A: B \stackrel{\text { def }}{=} \operatorname{tr}\left(A^{t} B\right)
$$

For $Q \in \mathbb{R}^{2 \times 2}$, we use $|Q|$ to denote its Frobenius norm, i.e.,

$$
|Q| \stackrel{\text { def }}{=} \sqrt{\operatorname{tr}\left(Q^{t} Q\right)}=\sqrt{\sum_{1 \leq i, j \leq 2} Q^{i j} Q^{i j}} .
$$

Besides, we define the matrix valued $L^{p}(1 \leq p \leq \infty)$ space by

$$
L^{p}\left(\Omega \rightarrow \mathbb{R}^{2 \times 2}\right) \stackrel{\text { def }}{=}\left\{Q: \Omega \rightarrow \mathbb{R}^{2 \times 2},|Q| \in L^{p}(\Omega, \mathbb{R})\right\}
$$

Further, for any smooth scalar function $u: \Omega \rightarrow \mathbb{R}$, we define the following Hölder norms and semi-norms:

$$
\begin{array}{ll}
{[u]_{C^{\alpha}(\bar{\Omega})} \stackrel{\text { def }}{=} \sup _{x \neq y \in \Omega} \frac{|u(x)-u(y)|}{|x-y|^{\alpha}},} & 0<\alpha \leq 1 . \\
{[u]_{C^{1+\alpha}(\bar{\Omega})} \stackrel{\text { def }}{=} \max _{1 \leq i \leq 2}\left[\partial_{i} u\right]_{C^{\alpha}(\bar{\Omega})},} & {[u]_{C^{2+\alpha}} \stackrel{\text { def }}{=} \max _{1 \leq i, j \leq 2}\left[\partial_{i} \partial_{j} u\right]_{C^{\alpha}(\bar{\Omega})},} \\
\|u\|_{C^{0}(\bar{\Omega})} \stackrel{\text { def }}{=} \sup _{x \in \Omega}|u(x)|, & 0<\alpha \leq 1 . \\
\|u\|_{C^{\alpha}(\bar{\Omega})} \stackrel{\text { def }}{=}\|u\|_{C^{0}(\Omega)}+[u]_{C^{\alpha}(\bar{\Omega})}, & 0<\alpha \leq 1 . \\
\|u\|_{C^{1+\alpha}(\bar{\Omega})} \stackrel{\text { def }}{=}\|u\|_{C^{1}(\bar{\Omega})}+[u]_{C^{1+\alpha}(\bar{\Omega})}, & 0<\alpha \leq 1 . \\
\|u\|_{C^{2+\alpha}(\bar{\Omega})} \stackrel{\text { def }}{=}\|u\|_{C^{2}(\bar{\Omega})}+[u]_{C^{2+\alpha}(\bar{\Omega})}, & 0<\alpha<1 .
\end{array}
$$

For a tensor valued function $Q: \Omega \rightarrow \mathbb{R}^{2 \times 2}$, the corresponding norms are defined to be the maximum of each component, for instance, $[Q]_{C^{\alpha}(\bar{\Omega})}=\max _{1 \leq i, j \leq 2}\left[Q^{i j}\right]_{C^{\alpha}(\bar{\Omega})}$; and the corresponding Hölder space by

$$
C^{\alpha}\left(\bar{\Omega} \rightarrow \mathbb{R}^{2 \times 2}\right) \stackrel{\text { def }}{=}\left\{Q: \Omega \rightarrow \mathbb{R}^{2 \times 2}, \max _{1 \leq i, j \leq 2}\left[Q^{i j}\right]_{\alpha} \in C^{\alpha}(\bar{\Omega})\right\}
$$

Without ambiguity, $L^{p}\left(\Omega \rightarrow \mathbb{R}^{2 \times 2}\right)$ will often be abbreviated as $L^{p}(\Omega)(1 \leq p \leq \infty)$, and $C^{k+\alpha}(\bar{\Omega} \rightarrow$ $\left.\mathbb{R}^{2 \times 2}\right)$ as $C^{k+\alpha}(\bar{\Omega})\left(0 \leq \alpha<1, k \in \mathbb{Z}^{+}\right)$. For the sake of simplicity, we at times use $\|\cdot\|_{L^{p}}$ to denote $\|\cdot\|_{L^{p}(\Omega)}$, and $\|\cdot\|_{C^{k+\alpha}}$ to denote $\|\cdot\|_{C^{k+\alpha}(\bar{\Omega})}$, respectively. We denote the partial derivative with respect to $x_{k}$ of the $i j$ component of $Q$, by $\partial_{k} Q^{i j}$.

## 2 Time discretization and its analysis

Let $\tilde{Q} \in C^{2+\alpha}(\bar{\Omega})$. We start with $Q^{0} \in C^{2+\alpha}(\bar{\Omega})$, and for $n=0,1,2, \cdots$, and $\Delta t>0$, find $Q^{n+1}$ from the following stabilized discretizations for (1.9)-(1.10):

$$
\frac{Q^{i j, n+1}-Q^{i j, n}}{\Delta t}=\zeta \Delta Q^{i j, n+1}-a Q^{i j, n+1}-c\left|Q^{n+1}\right|^{2} Q^{i j, n+1}-L\left(Q^{i j, n+1}-Q^{i j, n}\right)\left|\nabla Q^{n+1}\right|^{2}
$$

$$
\begin{align*}
& \quad+L_{4}\left\{2 \partial_{k}\left(Q^{l k, n} \partial_{l} Q^{i j, n+1}\right)-\partial_{i} Q^{k l, n+1} \partial_{j} Q^{k l, n+1}+\frac{\left|\nabla Q^{n+1}\right|^{2}}{2} \delta^{i j}\right\},  \tag{2.1}\\
& \left.Q^{n+1}\right|_{\partial \Omega}=\tilde{Q} ; \quad 1 \leq i, j \leq 2 . \tag{2.2}
\end{align*}
$$

Several remarks are in order:

- The above scheme is essentially a backward Euler scheme with an additional stabilizing term $-L\left(Q^{i j, n+1}-Q^{i j, n}\right)\left|\nabla Q^{n+1}\right|^{2}$ which plays an essential role in our analysis below. The stabilizing constant $L>0$ is to be determined later (cf. (2.4)).
- It is easy to see that (2.1) is a first order accurate approximation to (1.9).
- (2.1) can be simplified by taking into account of the traceless and symmetry properties of the Q-tensor function (cf. (3.1)), but we consider the current form (2.1) for generality.

Our main result regarding the convergence of (2.1)-(2.2) is stated in Theorem 2.2. Before proving the convergence, we are going to establish the unique solvability first, since the scheme (2.1)-(2.2) is highly nonlinear and its solvability is non-trivial.

### 2.1 A priori estimates and well-posedness of (2.1)-(2.2)

We start with some a priori estimates for the time-discrete problem (2.1)-(2.2).
Lemma 2.1. Let $Q^{n} \in C^{2+\alpha}(\bar{\Omega})$, and assume

$$
\begin{equation*}
\max \left\{\left\|Q^{n}\right\|_{L^{\infty}(\Omega)},\|\tilde{Q}\|_{L^{\infty}(\partial \Omega)}\right\} \leq \frac{\zeta}{(1+\sqrt{6})\left|L_{4}\right|} \tag{2.3}
\end{equation*}
$$

and that the stabilized constant $L$ satisfies

$$
\begin{equation*}
L \geq \frac{(\sqrt{6}+1)\left|L_{4}\right|^{2}}{\zeta} \tag{2.4}
\end{equation*}
$$

where $\zeta$ is defined in (1.11). Then, if $Q^{n+1} \in C^{2+\alpha}(\bar{\Omega})$ is a classical solution of (2.1)-(2.2), it holds

$$
\begin{equation*}
\left\|Q^{n+1}\right\|_{L^{\infty}(\Omega)} \leq \max \left\{\left\|Q^{n}\right\|_{L^{\infty}(\Omega)}, \sqrt{\frac{a^{-}}{c}}\right\} \tag{2.5}
\end{equation*}
$$

where $a^{-}=\max \{0,-a\}$.
Proof. Denoting

$$
\begin{equation*}
\rho^{n}=\left|Q^{n}\right|^{2}, \quad \rho^{n+1}=\left|Q^{n+1}\right|^{2} \tag{2.6}
\end{equation*}
$$

and multiplying both sides of (2.1) with $2 Q^{i j, n+1}$, then summing up for $1 \leq i, j \leq 2$, we get

$$
\begin{aligned}
& \frac{\left|Q^{n+1}\right|^{2}+\left|Q^{n+1}-Q^{n}\right|^{2}-\left|Q^{n}\right|^{2}}{\Delta t} \\
& \quad=\partial_{k}\left[\left(\zeta \delta^{k l}+2 L_{4} Q^{l k, n}\right) \partial_{l} \rho^{n+1}\right]-\left(4 L_{4} Q^{l k, n}+2 \zeta \delta^{k l}\right) \partial_{k} Q^{i j, n+1} \partial_{l} Q^{i j, n+1}
\end{aligned}
$$

$$
\begin{align*}
& -2 L_{4} Q^{i j, n+1} \partial_{i} Q^{k l, n+1} \partial_{j} Q^{k l, n+1}+L_{4} \operatorname{tr}\left(Q^{n+1}\right)\left|\nabla Q^{n+1}\right|^{2}-\left(a+c\left|Q^{n+1}\right|^{2}\right)\left|Q^{n+1}\right|^{2} \\
& -2 L\left[\left|Q^{n+1}\right|^{2}-\operatorname{tr}\left(\left(Q^{n}\right)^{t} Q^{n+1}\right)\right]\left|\nabla Q^{n+1}\right|^{2} . \tag{2.7}
\end{align*}
$$

Let us assume $\rho^{n+1}(\cdot)$ take its maximum value at some point $x_{0} \in \Omega$. Evaluating equation (2.1) at $x_{0}$, then we have
Case 1: if $\left|Q^{n+1}\right|\left(x_{0}\right) \leq \sqrt{\frac{a^{-}}{c}}$, then the proof is complete.
Case 2: otherwise, we can assume $\sqrt{\rho^{n+1}\left(x_{0}\right)}>\sqrt{\rho^{n}\left(x_{0}\right)}$, because $\sqrt{\rho^{n+1}\left(x_{0}\right)} \leq \sqrt{\rho^{n}\left(x_{0}\right)}$ will yield the conclusion (2.5) directly. First, for any matrix $Q \in \mathbb{R}^{2 \times 2}$ and row vector $b=\left(b^{1}, b^{2}\right)$, using Cauchy-Schwarz inequality, we have

$$
\left|Q^{i j} b^{i} b^{j}\right| \leq \frac{1}{2}\left(\left(2\left|Q^{11}\right|+\left|Q^{12}\right|+\left|Q^{21}\right|\right)\left|b^{1}\right|^{2}+\left(\left|Q^{12}\right|+\left|Q^{21}\right|+2\left|Q^{22}\right|\right)\left|b^{2}\right|^{2}\right) \leq \frac{\sqrt{6}}{2}|Q|\left(\left|b^{1}\right|^{2}+\left|b^{2}\right|^{2}\right) .
$$

As a consequence, it holds

$$
\begin{equation*}
-\left(4 L_{4} Q^{l k, n}+2 \zeta \delta^{k l}\right) \partial_{k} Q^{i j, n+1} \partial_{l} Q^{i j, n+1} \leq-\left(2 \zeta-2 \sqrt{6}\left|L_{4}\right| \sqrt{\rho^{n}}\right)\left|\nabla Q^{n+1}\right|^{2} . \tag{2.8}
\end{equation*}
$$

Besides, using Cauchy-Schwarz repeatedly we get

$$
\begin{align*}
& -2 L_{4} Q^{i j, n+1} \partial_{i} Q^{k l, n+1} \partial_{j} Q^{k l, n+1}+L_{4} \operatorname{tr}\left(Q^{n+1}\right)\left|\nabla Q^{n+1}\right|^{2} \\
& \quad=-2 L_{4}\left(Q^{12, n+1}+Q^{21, n+1}\right) \partial_{1} Q^{k l, n+1} \partial_{2} Q^{k l, n+1}-L_{4}\left(Q^{11, n+1}-Q^{22, n+1}\right) \partial_{1} Q^{k l, n+1} \partial_{1} Q^{k l, n+1} \\
& \quad \quad+L_{4}\left(Q^{11, n+1}-Q^{22, n+1}\right) \partial_{2} Q^{k l, n+1} \partial_{2} Q^{k l, n+1} \\
& \leq\left|L_{4}\right|\left(\left|Q^{11, n+1}\right|+\left|Q^{12, n+1}\right|+\left|Q^{21, n+1}\right|+\left|Q^{22, n+1}\right|\right)\left|\nabla Q^{n+1}\right|^{2} \\
& \leq \tag{2.9}
\end{align*}
$$

Using (2.8), (2.9) and the assumption (1.8), we see that (2.7) is reduced to

$$
\begin{align*}
\frac{\rho^{n+1}\left(x_{0}\right)-\rho^{n}\left(x_{0}\right)}{\Delta t} \leq & -\left(2 \zeta-2 \sqrt{6}\left|L_{4}\right| \sqrt{\rho^{n}}\right)\left|\nabla Q^{n+1}\right|^{2}+2\left|L_{4}\right| \sqrt{\rho^{n+1}}\left|\nabla Q^{n+1}\right|^{2} \\
& -c\left(\rho^{n+1}-\frac{a^{-}}{c}\right) \rho^{n+1}-2 L\left(\rho^{n+1}-\sqrt{\rho^{n}} \sqrt{\rho^{n+1}}\right)\left|\nabla Q^{n+1}\right|^{2}  \tag{2.10}\\
\leq- & {\left[2 \zeta-2 \sqrt{6}\left|L_{4}\right| \sqrt{\rho^{n}}+2 L \rho^{n+1}-\left(2\left|L_{4}\right|+2 L \sqrt{\rho^{n}}\right) \sqrt{\rho^{n+1}}\right]\left|\nabla Q^{n+1}\right|^{2} . }
\end{align*}
$$

Note that the quadratic function $2 L \rho^{n+1}-\left(2\left|L_{4}\right|+2 L \sqrt{\rho^{n}}\right) \sqrt{\rho^{n+1}}$ in terms of $\sqrt{\rho^{n+1}}$ in (2.10) is monotone increasing in the interval $I_{\rho^{n}}=\left[\frac{\left|L_{4}\right|}{2 L}+\frac{1}{2} \sqrt{\rho^{n}}, \infty\right)$, and attains its minimum at $\frac{\left|L_{4}\right|}{2 L}+\frac{1}{2} \sqrt{\rho^{n}}$.

If $\sqrt{\rho^{n}\left(x_{0}\right)} \geq \frac{\left|L_{4}\right|}{L}$, then $\sqrt{\rho^{n+1}\left(x_{0}\right)}>\sqrt{\rho^{n}\left(x_{0}\right)} \geq \frac{\left|L_{4}\right|}{2 L}+\frac{1}{2} \sqrt{\rho^{n}\left(x_{0}\right)}$, and based on (2.3), it is easy to check from equation (2.10) that for the case $\sqrt{\rho^{n}\left(x_{0}\right)}>\frac{\left|L_{4}\right|}{L}$,

$$
\frac{\rho^{n+1}\left(x_{0}\right)-\rho^{n}\left(x_{0}\right)}{\Delta t} \leq-\left[2 \zeta-2 \sqrt{6}\left|L_{4}\right| \sqrt{\rho^{n}\left(x_{0}\right)}-2\left|L_{4}\right| \sqrt{\rho^{n}\left(x_{0}\right)}\right]\left|\nabla Q^{n+1}\left(x_{0}\right)\right|^{2} \leq 0,
$$

which yields $\left|Q^{n+1}\right|\left(x_{0}\right) \leq\left\|Q^{n}\right\|_{L^{\infty}(\Omega)}$ so (2.5) holds.

On the other hand, if $\sqrt{\rho^{n}\left(x_{0}\right)}<\frac{\left|L_{4}\right|}{L}$, similarly as above, we derive from equation (2.10) that, for $L$ satisfying (2.4), it holds

$$
\begin{aligned}
\frac{\rho^{n+1}\left(x_{0}\right)-\rho^{n}\left(x_{0}\right)}{\Delta t} & \leq-\left[2 \zeta-2 \sqrt{6}\left|L_{4}\right| \sqrt{\rho^{n}\left(x_{0}\right)}-\frac{L}{2}\left(\frac{\left|L_{4}\right|}{L}+\sqrt{\rho^{n}\left(x_{0}\right)}\right)^{2}\right]\left|\nabla Q^{n+1}\left(x_{0}\right)\right|^{2} \\
& \leq-\left[2 \zeta-2 \sqrt{6}\left|L_{4}\right| \frac{\zeta}{(1+\sqrt{6})\left|L_{4}\right|}-\frac{L}{2}\left(\frac{\left|L_{4}\right|}{L}+\frac{\left|L_{4}\right|}{L}\right)^{2}\right]\left|\nabla Q^{n+1}\left(x_{0}\right)\right|^{2} \\
& \leq-\left[2 \frac{\zeta}{(1+\sqrt{6})}-\frac{2\left|L_{4}\right|^{2}}{L}\right]\left|\nabla Q^{n+1}\left(x_{0}\right)\right|^{2} \\
& \leq 0,
\end{aligned}
$$

which immediately implies (2.5).
Combining all the arguments above, the proof is complete.
Remark 2.1. It is easy to check from the proof of Lemma 2.1 that (2.5) is still valid if the R.H.S. of (2.3) is replaced by any sufficiently small constant $\eta>0$, and $L$ satisfies (2.4).

Note that in the above Lemma 2.1 we proved that for classical solutions, the $L^{\infty}$ norm at the ( $n+1$ )-th step will remain to be small, provided that the $L^{\infty}$ norm at the $n$-th step is assumed to be small (small boundary data and $\frac{a^{-}}{c}$ as well). But we have not yet proved the existence of such classical solutions to (2.1)-(2.2). To this end, we shall apply the Leray-Schauder theory for the existence of classical solutions. For the reader's convenience, first we recall below the LeraySchauder fixed point theorem [13].

Theorem 2.1 (Leray-Schauder fixed point theorem). Let $\mathcal{B}$ be a Banach space and $\mathcal{T}: \mathcal{B} \times[0,1] \rightarrow$ $\mathcal{B}$ a compact map such that

1. $\mathcal{T}(x, 0)=0, \forall x \in \mathcal{B}$,
2. there exists a constant $M>0$ such that for each pair $(x, \sigma) \in \mathcal{B} \times[0,1]$ which satisfies $x=\mathcal{T}(x, \sigma)$, we have

$$
\begin{equation*}
\|x\|<M . \tag{2.11}
\end{equation*}
$$

Then the map $\mathcal{T}_{1}: \mathcal{B} \rightarrow \mathcal{B}$ given by $\mathcal{T}_{1} y=\mathcal{T}(y, 1), y \in \mathcal{B}$ has a fixed point.
By virtue of Theorem 2.1, we have
Proposition 2.1. Let $Q^{n} \in C^{2+\alpha}\left(\bar{\Omega} \rightarrow \mathbb{R}^{2 \times 2}\right)$. Suppose $\left\|Q^{n}\right\|_{C^{0}(\bar{\Omega})},\|\tilde{Q}\|_{C^{0}(\partial \Omega)}$ and $\frac{a^{-}}{c}$ are sufficiently small and L satisfies (2.4). Then there exists a classical solution $Q^{n+1} \in C^{2+\alpha}\left(\bar{\Omega} \rightarrow \mathbb{R}^{2 \times 2}\right)$ to the system (2.1)-(2.2). Furthermore, (2.5) is also satisfied.

Proof. To utilize Theorem 2.1, we define

$$
\mathcal{B}=C^{1+\alpha}\left(\bar{\Omega} \rightarrow \mathbb{R}^{2 \times 2}\right),
$$

and a map

$$
\mathcal{T}: \mathcal{B} \times[0,1] \rightarrow \mathcal{B}
$$

Here $\mathcal{T}(u, \theta) \in C^{2+\alpha}\left(\bar{\Omega} \rightarrow \mathbb{R}^{2 \times 2}\right) \subset \mathcal{B}$ with $u \in \mathcal{B}, \theta \in[0,1]$ solves the equation

$$
\begin{aligned}
& \theta\left\{\zeta \Delta w^{i j}+2 L_{4} \partial_{k}\left(Q^{l k, n} \partial_{l} w^{i j}\right)-L_{4} \partial_{i} u^{k l} \partial_{j} u^{k l}+\frac{L_{4}}{2}|\nabla u|^{2} \delta^{i j}-\left(a+c|u|^{2}\right) u^{i j}\right. \\
& \left.\quad-L\left(u^{i j}-Q^{i j, n}\right)|\nabla u|^{2}-\frac{u^{i j}-Q^{i j, n}}{\Delta t}\right\}+(1-\theta) \Delta w^{i j}=0, \quad 1 \leq i, j \leq 2, \\
& \left.w\right|_{\partial \Omega}=\theta \tilde{Q} .
\end{aligned}
$$

We proceed to prove that all conditions in Theorem 2.1 are satisfied. To begin with, it is easy to see that $\mathcal{T}(u, 0)=0, \forall u \in \mathcal{B}$. Next we assume $(u, \sigma) \in \mathcal{B} \times[0,1]$ satisfies $u=\mathcal{T}(u, \sigma)$, that is,

$$
\begin{align*}
\partial_{k}\{ & {\left.\left[(\sigma \zeta+1-\sigma) \delta^{k l}+2 \sigma L_{4} Q^{n}\right] \partial_{l} u^{i j}\right\} } \\
& =\sigma\left\{L_{4} \partial_{i} u^{k l} \partial_{j} u^{k l}-\frac{L_{4}}{2}|\nabla u|^{2} \delta^{i j}+\left(a+c|u|^{2}\right) u^{i j}+L\left(u^{i j}-Q^{i j, n}\right)|\nabla u|^{2}+\frac{u^{i j}-Q^{i j, n}}{\Delta t}\right\} \\
& \doteq \sigma f^{i j}, \quad 1 \leq i, j \leq 2  \tag{2.12}\\
\left.u\right|_{\partial \Omega} & =\sigma \tilde{Q} . \tag{2.13}
\end{align*}
$$

Then, following the same procedure in Lemma 2.1, one may conclude

$$
\begin{equation*}
\|u\|_{C^{0}} \leq \max \left\{\left\|Q^{n}\right\|_{C^{0}}, \sqrt{\frac{a^{-}}{c}}\right\} \tag{2.14}
\end{equation*}
$$

provided that $\left\|Q^{n}\right\|_{C^{0}},\|\tilde{Q}\|_{C^{0}(\partial \Omega)}$ and $\frac{a^{-}}{c}$ are sufficiently small. As a consequence, using the classical Schauder estimate (see Theorem 6.6 in [13]), interpolation inequality and Young's inequality, one can derive from (2.12)-(2.13) that for sufficiently small $\left\|Q^{n}\right\|_{C^{0}}$, we have

$$
\begin{align*}
\|u\|_{C^{2+\alpha}} \leq & C\|u\|_{C^{0}}+C\|\tilde{Q}\|_{C^{2+\alpha}}+C\|\sigma f\|_{C^{\alpha}} \\
\leq & C\left\|Q^{0}\right\|_{C^{0}}+C+C\|f\|_{C^{\alpha}} \\
\leq & C+C\left\||\nabla u|^{2}\right\|_{C^{\alpha}}+C\left\|a u+c|u|^{2} u\right\|_{C^{\alpha}}+C\left\|u|\nabla u|^{2}\right\|_{C^{\alpha}} \\
& +C\left\|Q^{n}|\nabla u|^{2}\right\|_{C^{\alpha}}+C\left(\|u\|_{C^{\alpha}}+\left\|Q^{n}\right\|_{C^{\alpha}}\right) \\
\leq & C+C\|u\|_{C^{\alpha}}+C\left\||\nabla u|^{2}\right\|_{C^{\alpha}}+C\left\|u|\nabla u|^{2}\right\|_{C^{\alpha}}+C\left\|Q^{n}|\nabla u|^{2}\right\|_{C^{\alpha}} \\
\leq & C+C\|u\|_{C^{0}}^{\frac{2}{2+\alpha}}\|u\|_{C^{2}+\alpha}^{\frac{\alpha}{2+\alpha}}+C\left\|\left|\nabla u\left\|_{C^{0}}\right\| \nabla u u\left\|_{C^{\alpha}}+C\left(\|u\|_{C^{0}}+\left\|Q^{n}\right\|_{C^{0}}\right)\right\| \nabla u\right|^{2}\right\|_{C^{\alpha}} \\
& +C\left(\|u\|_{C^{\alpha}}+\left\|Q^{n}\right\|_{C^{\alpha}}\right)\left\||\nabla u|^{2}\right\|_{C^{0}} \\
\leq & C+C\|u\|_{C^{0}}^{\frac{2}{2+\alpha}}\|u\|_{C^{2+\alpha}}^{\frac{\alpha}{2+\alpha}}+C\|u\|_{C^{0}}\|u\|_{C^{2+\alpha}}+C\left(\|u\|_{C^{0}}+\left\|Q^{n}\right\|_{C^{0}}\right)\|u\|_{C^{0}}\|u\|_{C^{2+\alpha}} \\
& +C\|u\|_{C^{0}}^{2}\|u\|_{C^{2+\alpha}}+C\left\|Q^{n}\right\|_{C^{0}}^{\frac{2}{2+\alpha}}\left\|Q^{n}\right\|_{C^{2+\alpha}}^{\frac{\alpha}{2+\alpha}}\|u\|_{C^{0}}^{\frac{2+2 \alpha}{2+\alpha}}\|u\|_{C^{2+\alpha}}^{\frac{2}{2+\alpha}} \\
\leq & C+\frac{1}{2}\|u\|_{C^{2+\alpha}} . \tag{2.15}
\end{align*}
$$

In the above $C>0$ is a generic constant that may depend on $\Omega, \Delta t,\left\|Q_{0}\right\|_{C^{0}},\|\tilde{Q}\|_{C^{2+\alpha}},\left\|Q^{n}\right\|_{C^{2+\alpha}}$ and coefficients of the system. Therefore (2.11) is valid. In addition, it is easy to check that $\mathcal{T}$ is a compact map due to the compact embedding $C^{2+\alpha}(\bar{\Omega}) \hookrightarrow C^{1+\alpha}(\bar{\Omega})$. Thus all conditions in Theorem 2.1 are satisfied and in conclusion $\mathcal{T}_{1} y=\mathcal{T}(y, 1)$ has a fixed point, which is equivalent to say that the system (2.1)-(2.2) admits a classical solution $Q^{n+1} \in C^{2+\alpha}(\bar{\Omega})$.

For classical solutions whose existence was proved in Proposition 2.1 above, we proceed to establish the following uniform estimates.

Proposition 2.2. The classical solutions established in Proposition 2.1 satisfy the following uniform bounds:

$$
\begin{align*}
& \quad\left\|\nabla Q^{n+1}\right\|_{L^{2}(\Omega)}^{2} \leq C \Delta t+C\left\|\Delta Q^{0}\right\|_{L^{2}(\Omega)}^{2} \Delta t+\left\|\nabla Q^{0}\right\|_{L^{2}(\Omega)}^{2}, \quad \forall 0 \leq n<\left[\frac{T}{\Delta t}\right]  \tag{2.16}\\
& \sum_{n=1}^{\left[\frac{T}{\Delta t}\right]}\left\|\Delta Q^{n}\right\|_{L^{2}(\Omega)}^{2} \Delta t \leq C T+\left\|\nabla Q^{0}\right\|_{L^{2}(\Omega)}^{2}+\left\|\Delta Q^{0}\right\|_{L^{2}(\Omega)}^{2} \Delta t, \tag{2.17}
\end{align*}
$$

provided that there exists a sufficiently small (but computable) constant $\varepsilon>0$ such that

$$
\begin{equation*}
\max \left\{\left\|Q^{0}\right\|_{L^{\infty}(\Omega)},\|\tilde{Q}\|_{L^{\infty}(\partial \Omega)}, \sqrt{\frac{a^{-}}{c}}\right\} \leq \varepsilon \tag{2.18}
\end{equation*}
$$

Here $C>0$ is a constant that only depends on $\zeta, \varepsilon, \Omega, a, c$ and $L_{4}$, but independent of $n$ or $\Delta t$.
Proof. Multiplying equation (2.1) with $-\Delta Q^{i j, n+1}$ and integrating over $\Omega$, we find

$$
\begin{align*}
& \frac{1}{2 \Delta t} \int_{\Omega}\left|\nabla Q^{n+1}\right|^{2}+\left|\nabla Q^{n+1}-\nabla Q^{n}\right|^{2}-\left|\nabla Q^{n}\right|^{2} d x \\
& =-\zeta \int_{\Omega}\left|\Delta Q^{n+1}\right|^{2} d x+a \int_{\Omega} Q^{i j, n+1} \Delta Q^{i j, n+1} d x+c \int_{\Omega}\left|Q^{n+1}\right|^{2} Q^{i j, n+1} \Delta Q^{i j, n+1} d x \\
& \quad+L \int_{\Omega}\left|\nabla Q^{n+1}\right|^{2}\left(Q^{i j, n+1}-Q^{i j, n}\right) \Delta Q^{i j, n+1} d x-2 L_{4} \int_{\Omega} Q^{l k, n} \partial_{k} \partial_{l} Q^{i j, n+1} \Delta Q^{i j, n+1} d x \\
& \quad-L_{4} \int_{\Omega}\left\{2 \partial_{k} Q^{l k, n} \partial_{l} Q^{i j, n+1}-\partial_{i} Q^{k l, n+1} \partial_{j} Q^{k l, n+1}+\frac{\left|\nabla Q^{n+1}\right|^{2}}{2} \delta^{i j}\right\} \Delta Q^{i j, n+1} d x \\
& =\zeta \int_{\Omega}\left|\Delta Q^{n+1}\right|^{2} d x+I_{1}+\cdots+I_{5} . \tag{2.19}
\end{align*}
$$

We estimate below the terms $I_{1}$ through $I_{5}$ individually. To begin with, it follows from (2.5) and (2.18) that

$$
\begin{equation*}
\left\|Q^{n+1}\right\|_{L^{\infty}(\Omega)} \leq \varepsilon \tag{2.20}
\end{equation*}
$$

Using Young's inequality and (2.20), we obtain

$$
\begin{aligned}
I_{1}+I_{2} & \leq \frac{\zeta}{8}\left\|\Delta Q^{n+1}\right\|^{2}+C\left(\left\|Q^{n+1}\right\|_{L^{2}}^{2}+\left\|Q^{n+1}\right\|_{L^{6}}^{6}\right) \\
& \leq \frac{\zeta}{8}\left\|\Delta Q^{n+1}\right\|^{2}+C\left(\varepsilon^{2}+\varepsilon^{6}\right)|\Omega|
\end{aligned}
$$

$$
\begin{equation*}
\leq \frac{\zeta}{8}\left\|\Delta Q^{n+1}\right\|^{2}+C \tag{2.21}
\end{equation*}
$$

By (2.20), Gagliardo-Nirenberg interpolation inequality and classical elliptic PDE theory [10], we get

$$
\begin{align*}
I_{3} & \leq L\left\|\nabla Q^{n+1}\right\|_{L^{4}}^{2}\left(\left\|Q^{n+1}\right\|_{L^{\infty}}+\left\|Q^{n}\right\|_{L^{\infty}}\right)\left\|\Delta Q^{n+1}\right\|_{L^{2}} \\
& \leq C\left(\left\|Q^{n+1}\right\|_{L^{\infty}}+\left\|Q^{n}\right\|_{L^{\infty}}\right)\left\|Q^{n+1}\right\|_{L^{\infty}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|Q^{n+1}\right\|_{L^{2}}+\|\tilde{Q}\|_{H^{\frac{3}{2}}(\partial \Omega)}\right)\left\|\Delta Q^{n+1}\right\|_{L^{2}} \\
& \leq C \varepsilon^{2}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|Q^{n+1}\right\|_{L^{2}}+\|\tilde{Q}\|_{H^{\frac{3}{2}}(\partial \Omega)}\right)\left\|\Delta Q^{n+1}\right\|_{L^{2}} \\
& \leq \frac{\zeta}{8}\left\|\Delta Q^{n+1}\right\|^{2}+C \tag{2.22}
\end{align*}
$$

and

$$
\begin{align*}
I_{4} & \leq 2\left|L_{4}\right|\left\|Q^{n}\right\|_{L^{\infty}}\left\|\nabla^{2} Q^{n+1}\right\|_{L^{2}}\left\|\Delta Q^{n+1}\right\|_{L^{2}} \\
& \leq C\left\|Q^{n}\right\|_{L^{\infty}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|Q^{n+1}\right\|_{L^{2}}+\|\tilde{Q}\|_{H^{\frac{3}{2}}(\partial \Omega)}\right)\left\|\Delta Q^{n+1}\right\|_{L^{2}} \\
& \leq \frac{\zeta}{8}\left\|\Delta Q^{n+1}\right\|^{2}+C . \tag{2.23}
\end{align*}
$$

Similarly as in the estimate of $I_{3}$, we can obtain

$$
\begin{align*}
I_{5} & \leq 2\left|L_{4}\right|\left\|\nabla Q^{n}\right\|_{L^{4}}\left\|\nabla Q^{n+1}\right\|_{L^{4}}\left\|\Delta Q^{n+1}\right\|_{L^{2}}+2\left|L_{4}\right|\left\|\nabla Q^{n+1}\right\|_{L^{4}}^{2}\left\|\Delta Q^{n+1}\right\|_{L^{2}} \\
& \leq \frac{\zeta}{4}\left\|\Delta Q^{n}\right\|^{2}+\frac{\zeta}{8}\left\|\Delta Q^{n+1}\right\|^{2}+C . \tag{2.24}
\end{align*}
$$

Combining the above we conclude that $\forall 0 \leq n<\left[\frac{T}{\Delta t}\right]$, it holds

$$
\begin{equation*}
\left\|\nabla Q^{n+1}\right\|_{L^{2}}^{2}+\left\|\nabla Q^{n+1}-\nabla Q^{n}\right\|_{L^{2}}^{2}-\left\|\nabla Q^{n}\right\|_{L^{2}}^{2} \leq-\zeta\left\|\Delta Q^{n+1}\right\|_{L^{2}}^{2} \Delta t+\frac{\zeta}{2}\left\|\Delta Q^{n}\right\|_{L^{2}}^{2} \Delta t+C \Delta t \tag{2.25}
\end{equation*}
$$

As a consequence, summing up the above estimate (2.25) for $n$ from 0 to $\left[\frac{T}{\Delta t}\right]-1$ leads to (2.16) and (2.17).

Based on the uniform estimates (2.16) and (2.17) established in Proposition 2.2, we can further obtain the uniqueness result concerning the classical solutions of the system (2.1)-(2.2).

Proposition 2.3. Let $Q^{n} \in C^{2+\alpha}(\bar{\Omega})$. Suppose $P^{n+1}, Q^{n+1} \in C^{2+\alpha}(\bar{\Omega})$ are two classical solutions to the problem (2.1)-(2.2) that satisfy (2.20). If $\varepsilon$ in Proposition 2.2 is chosen to be suitably small (but independent of $n$ or $\Delta t$ ), then

$$
P^{n+1} \equiv Q^{n+1}
$$

Proof. Let $\bar{R}^{n+1}=Q^{n+1}-P^{n+1}$. We have

$$
\frac{\bar{R}^{i j, n+1}}{\Delta t}=\zeta \Delta \bar{R}^{i j, n+1}-a \bar{R}^{i j, n+1}-c\left(\left|Q^{n+1}\right|^{2} Q^{i j, n+1}-\left|P^{n+1}\right|^{2} P^{i j, n+1}\right)
$$

$$
\begin{align*}
&-L\left(Q^{i j, n+1}\left|\nabla Q^{n+1}\right|^{2}-P^{i j, n+1}\left|\nabla P^{n+1}\right|^{2}\right)+L Q^{i j, n}\left(\left|\nabla Q^{n+1}\right|^{2}-\left|\nabla P^{n+1}\right|^{2}\right) \\
&+2 L_{4} \partial_{k}\left(Q^{l k, n} \partial_{l} \bar{R}^{i j, n+1}\right)-L_{4}\left(\partial_{i} Q^{l k, n+1} \partial_{j} Q^{l k, n+1}-\partial_{i} P^{l k, n+1} \partial_{j} P^{l k, n+1}\right) \\
&+\frac{L_{4}}{2}\left(\left|\nabla Q^{n+1}\right|^{2}-\left|\nabla P^{n+1}\right|^{2}\right) \delta^{i j}  \tag{2.26}\\
&\left.\bar{R}^{n+1}\right|_{\partial \Omega}=0 \tag{2.27}
\end{align*}
$$

Multiplying equation (2.26) with $\bar{R}^{n+1}$, integrating over $\Omega$ and using the boundary condition (2.27), we obtain

$$
\begin{align*}
& \frac{\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}}{\Delta t} \\
& =-\zeta\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}-a\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}-c \int_{\Omega}\left(\left|Q^{n+1}\right|^{2} Q^{i j, n+1}-\left|P^{n+1}\right|^{2} P^{i j, n+1}\right) \bar{R}^{i j, n+1} d x \\
& \quad-L \int_{\Omega}\left(\left|\nabla Q^{n+1}\right|^{2} Q^{i j, n+1}-\left|\nabla P^{n+1}\right|^{2} P^{i j, n+1}\right) \bar{R}^{i j, n+1} d x \\
& \quad+L \int_{\Omega}\left(\left|\nabla Q^{n+1}\right|^{2}-\left|\nabla P^{n+1}\right|^{2}\right) Q^{i j, n} \bar{R}^{i j, n+1} d x-2 L_{4} \int_{\Omega} Q^{l k, n} \partial_{l} \bar{R}^{i j, n+1} \partial_{k} \bar{R}^{i j, n+1} d x \\
& \quad-L_{4} \int_{\Omega}\left(\partial_{i} Q^{l k, n+1} \partial_{j} Q^{l k, n+1}-\partial_{i} P^{l k, n+1} \partial_{j} P^{l k, n+1}\right) \bar{R}^{i j, n+1} d x \\
& \quad+\frac{L_{4}}{2} \int_{\Omega}\left(\left|\nabla Q^{n+1}\right|^{2}-\left|\nabla P^{n+1}\right|^{2}\right) \operatorname{tr}\left(\bar{R}^{n+1}\right) d x \\
& =-\zeta\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}+I_{1}+\cdots+I_{7} \tag{2.28}
\end{align*}
$$

Note that both $P^{n+1}$ and $Q^{n+1}$ satisfy (2.16)-(2.17) and (2.20). Hence

$$
\begin{aligned}
I_{1}+I_{2} & \leq-a\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}+c\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}\left(\left\|Q^{n+1}\right\|_{L^{\infty}}^{2}+\left\|Q^{n+1}\right\|_{L^{\infty}}\left\|P^{n+1}\right\|_{L^{\infty}}+\left\|P^{n+1}\right\|_{L^{\infty}}^{2}\right) \\
& \leq c\left(3 \varepsilon^{2}-\frac{a}{c}\right)\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2} \\
& \leq 4 c \varepsilon^{2}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}
\end{aligned}
$$

where we used (2.18) to derive the last inequality.
By using (2.5), Ladyzhenskaya's inequality, Gagliardo-Nirenberg inequality and classical elliptic PDE theory, we have

$$
\begin{aligned}
I_{3} \leq & L\left[\left\|\bar{R}^{n+1}\right\|_{L^{4}}^{2}\left\|\nabla Q^{n+1}\right\|_{L^{4}}^{2}+\left\|P^{n+1}\right\|_{L^{\infty}}\left(\left\|\nabla Q^{n+1}\right\|_{L^{4}}+\left\|\nabla P^{n+1}\right\|_{L^{4}}\right)\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}\left\|\bar{R}^{n+1}\right\|_{L^{4}}\right] \\
\leq & C L\left\|\bar{R}^{n+1}\right\|_{L^{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}\left\|Q^{n+1}\right\|_{L^{\infty}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|Q^{n+1}\right\|_{L^{2}}+\|\tilde{Q}\|_{H^{\frac{3}{2}}(\partial \Omega)}\right) \\
& +C L\left\|P^{n+1}\right\|_{L^{\infty}}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{\frac{1}{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{\frac{3}{2}}\left\|Q^{n+1}\right\|_{L^{\infty}}^{\frac{1}{2}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|Q^{n+1}\right\|_{L^{2}}+\|\tilde{Q}\|_{H^{\frac{3}{2}}(\partial \Omega)}\right)^{\frac{1}{2}} \\
& +C L\left\|P^{n+1}\right\|_{L^{\infty}}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{\frac{1}{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{\frac{3}{2}}\left\|P^{n+1}\right\|_{L^{\infty}}^{\frac{1}{2}}\left(\left\|\Delta P^{n+1}\right\|_{L^{2}}+\left\|P^{n+1}\right\|_{L^{2}}+\|\tilde{Q}\|_{H^{\frac{3}{2}}(\partial \Omega)}\right)^{\frac{1}{2}} \\
\leq & C L\left\|\bar{R}^{n+1}\right\|_{L^{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}\left\|Q^{0}\right\|_{L^{\infty}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+1\right) \\
& +C L\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{\frac{1}{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{\frac{3}{2}}\left\|Q^{0}\right\|_{L^{\infty}}^{\frac{3}{2}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|\Delta P^{n+1}\right\|_{L^{2}}+1\right)^{\frac{1}{2}}\left\|\bar{R}^{n+1}\right\|_{L^{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+1\right)
\end{aligned}
$$

$$
\begin{aligned}
&+C \varepsilon^{\frac{3}{2}}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{\frac{1}{2}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{\frac{3}{2}}\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}+\left\|\Delta P^{n+1}\right\|_{L^{2}}+1\right)^{\frac{1}{2}} \\
& \leq \frac{\zeta}{8}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}+C \varepsilon\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}^{2}+\left\|\Delta P^{n+1}\right\|_{L^{2}}^{2}+1\right)\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2} .
\end{aligned}
$$

Similarly,

$$
\begin{aligned}
I_{4} & \leq L\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}\left\|\bar{R}^{n+1}\right\|_{L^{4}}\left(\left\|\nabla Q^{n+1}\right\|_{L^{4}}+\left\|\nabla P^{n+1}\right\|_{L^{4}}\right)\left\|Q^{n}\right\|_{L^{\infty}} \\
& \leq \frac{\zeta}{8}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}+C \varepsilon\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}^{2}+\left\|\Delta P^{n+1}\right\|_{L^{2}}^{2}+1\right)\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}
\end{aligned}
$$

We derive from (2.5) that

$$
I_{5} \leq 2\left|L_{4}\right|\left\|Q^{n}\right\|_{L^{\infty}}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2} \leq 2\left|L_{4}\right| \varepsilon\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2} \leq \frac{\zeta}{8}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}
$$

We can control $I_{6}$ and $I_{7}$ in a manner similar for $I_{3}$, namely:

$$
\begin{aligned}
I_{6}+I_{7} & \leq 2\left|L_{4}\right|\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}\left\|\bar{R}^{n+1}\right\|_{L^{4}}\left(\left\|\nabla Q^{n+1}\right\|_{L^{4}}+\left\|\nabla P^{n+1}\right\|_{L^{4}}\right) \\
& \leq \frac{\zeta}{8}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}+C \varepsilon\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}^{2}+\left\|\Delta P^{n+1}\right\|_{L^{2}}^{2}+1\right)\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2} .
\end{aligned}
$$

After summing up the above inequalities in (2.28), we get

$$
\begin{equation*}
\frac{\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}}{\Delta t} \leq-\frac{\zeta}{2}\left\|\nabla \bar{R}^{n+1}\right\|_{L^{2}}^{2}+4 c \varepsilon^{2}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}+C \varepsilon\left(\left\|\Delta Q^{n+1}\right\|_{L^{2}}^{2}+\left\|\Delta P^{n+1}\right\|_{L^{2}}^{2}+1\right)\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2} \tag{2.29}
\end{equation*}
$$

Finally, we derive from (2.17) and the above inequality that

$$
\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2} \leq 4 c \varepsilon^{2}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}+C \varepsilon(2 C T+\Delta t)\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2} \leq \frac{1}{2}\left\|\bar{R}^{n+1}\right\|_{L^{2}}^{2}
$$

provided $\varepsilon$ is chosen to be sufficiently small. Therefore we conclude

$$
\bar{R}^{n+1} \equiv 0 .
$$

### 2.2 Convergence

Next we shall construct a family of approximate solutions using linear interpolation in time. The above a priori estimates for the set of discrete solutions allow us to obtain the existence of a time-continuous limit function which we will show to be a solution of the original PDE system (1.9)-(1.10).

Let us fix the initial data $Q^{0}$ and step size $h \stackrel{\text { def }}{=} \Delta t$ and define a piecewise linear interpolation $t \in[0, T) \rightarrow Q_{h}(\cdot, t)$ as

$$
\begin{equation*}
Q_{h}(x, t)=Q^{n}(x)+\frac{Q^{n+1}(x)-Q^{n}(x)}{h}(t-n h), \quad \forall x \in \Omega, \quad n h \leq t<(n+1) h, \quad 0 \leq n<\left[\frac{T}{h}\right] \tag{2.30}
\end{equation*}
$$

Based on equation (2.1) and the above construction (2.30), we know that $Q_{h}$ satisfies

$$
\begin{align*}
& \partial_{t} Q_{h}^{i j}(t, x)=\zeta \Delta Q_{h}^{i j}(x, n h)-a Q_{h}^{i j}(x, n h) \\
& -c\left|Q_{h}(x, n h)\right|^{2} Q_{h}^{i j}(x, n h)-L\left[Q_{h}^{i j}(n h, x)-Q_{h}^{i j}(n h-h, x)\right]\left|\nabla Q_{h}(n h, x)\right|^{2} \\
& \quad+L_{4}\left\{2 \partial_{k}\left[Q_{h}^{l k}(n h-h, x) \partial_{l} Q_{h}^{i j}(x, n h)\right]-\partial_{i} Q_{h}^{k l}(x, n h) \partial_{j} Q_{h}^{k l}(x, n h)+\frac{\left|\nabla Q_{h}(x, n h)\right|^{2}}{2} \delta^{i j}\right\} \\
& \quad \forall x \in \Omega, \quad(n-1) h \leq t<n h, \quad 1 \leq n \leq\left[\frac{T}{h}\right] \tag{2.31}
\end{align*}
$$

We collect from Proposition 2.2 and equation (2.31) the following uniform bounds

$$
\begin{align*}
& \quad\left\|\nabla Q_{h}(\cdot, t)\right\|_{L^{2}(\Omega)}^{2} \leq C \Delta t+C\left\|\Delta Q^{0}\right\|_{L^{2}(\Omega)}^{2} \Delta t+\left\|\nabla Q^{0}\right\|_{L^{2}(\Omega)}^{2}, \quad \forall t \in(0, T),  \tag{2.32}\\
& \int_{0}^{T}\left\|\Delta Q_{h}(\cdot, t)\right\|_{L^{2}(\Omega)}^{2} d t \leq C T+C\left\|\Delta Q^{0}\right\|_{L^{2}(\Omega)}^{2} \Delta t+\left\|\nabla Q^{0}\right\|_{L^{2}(\Omega)}^{2},  \tag{2.33}\\
& \int_{0}^{T}\left\|\partial_{t} Q_{h}(\cdot, t)\right\|_{L^{2}(\Omega)}^{2} d t \leq C T+C \tag{2.34}
\end{align*}
$$

In the above $C>0$ is a generic constant that does not depend on $h$.
As a consequence, as $h \rightarrow 0$, we have from Aubin-Lions Lemma (see [28]) that the following results hold.

Theorem 2.2. (Main result) Let $Q^{0}, \tilde{Q} \in C^{2+\alpha}\left(\bar{\Omega} \rightarrow \mathbb{R}^{2 \times 2}\right)$. Suppose $\left\|Q^{0}\right\|_{C^{0}(\bar{\Omega})},\|\tilde{Q}\|_{C^{0}(\partial \Omega)}$ and $\frac{a^{-}}{c}$ are sufficiently small and $L$ satisfies (2.4). Then the numerical scheme (2.1)-(2.2) admits unique solutions $Q^{n}$ for $n \geq 1$, and the piecewise linear interpolation $Q_{h}(t)$ of the numerical solution given in (2.30) converges to an exact solution of (1.9)-(1.10), i.e.

$$
\begin{aligned}
& Q_{h}(\cdot, t) \rightarrow Q(\cdot, t) \text { strong in } L^{2}\left(0, T ; H^{1}(\Omega)\right), \\
& Q_{h}(\cdot, t) \rightarrow Q(\cdot, t) \text { weakly in } L^{\infty}\left(0, T ; H^{1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{2}(\Omega)\right),
\end{aligned}
$$

where $Q(\cdot, t)$ solves

$$
\begin{aligned}
\partial_{t} Q^{i j} & =\zeta \Delta Q^{i j}-\left[a+c \operatorname{tr}\left(Q^{2}\right)\right] Q^{i j}+L_{4}\left\{2 \partial_{k}\left(Q^{l k} \partial_{l} Q^{i j}\right)-\partial_{i} Q^{k l} \partial_{j} Q^{k l}+\frac{|\nabla Q|^{2} \delta^{i j}}{2}\right\}, \\
\left.Q\right|_{\partial \Omega} & =\tilde{Q}, \quad Q(0, x)=Q^{0}(x)
\end{aligned}
$$

in the weak sense defined in Definition 2.1 below.
We can also check directly that the limit solution $Q$ always lies in the $Q$-tensor space $\mathcal{S}^{(2)}$, provided $Q^{0}, \tilde{Q} \in \mathcal{S}^{(2)}$.

Next, we recall the notion of weak solutions discussed in [15]
Definition 2.1. For any $T \in(0,+\infty)$, a function $Q$ satisfying
$Q \in L^{\infty}\left(0, T ; H^{1} \cap L^{\infty}\right) \cap L^{2}\left(0, T ; H^{2}\right), \quad \partial_{t} Q \in L^{2}\left(0, T ; L^{2}\right), \quad$ and $\quad Q \in S^{(2)}, \quad$ a.e. in $\Omega \times(0, T)$,
is called a weak solution of the problem (1.9)-(1.10), if it satisfies the initial and boundary conditions (1.10), and we have

$$
\begin{aligned}
-\int_{\Omega \times[0, T]} Q: \partial_{t} R d x d t=- & 2 L_{1} \int_{\Omega \times[0, T]} \partial_{k} Q: \partial_{k} R d x d t-\int_{\Omega \times[0, T]}\left[a+c \operatorname{tr}\left(Q^{2}\right)\right] Q: R d x d t \\
& -2\left(L_{2}+L_{3}\right) \int_{\Omega \times[0, T]} \partial_{k} Q_{i k} \partial_{j} R_{i j} d x d t+\left(L_{2}+L_{3}\right) \int_{\Omega \times[0, T]} \partial_{k} Q_{l k} \partial_{l} R_{i i} d x d t \\
& -2 L_{4} \int_{\Omega \times[0, T]} Q_{l k} \partial_{k} Q_{i j} \partial_{l} R_{i j} d x d t-L_{4} \int_{\Omega \times[0, T]} \partial_{i} Q_{k l} \partial_{j} Q_{k l} R_{i j} d x d t \\
& +\frac{L_{4}}{2} \int_{\Omega \times[0, T]}|\nabla Q|^{2} R_{i i} d x d t-\int_{\Omega} Q_{0}: R(0) d x .
\end{aligned}
$$

Here $R \in C_{c}^{\infty}\left([0, T) \times \Omega \rightarrow \mathbb{R}^{2 \times 2}\right)$ is arbitrary.
Summing up the above, we obtained the well-posedness result for (1.9)-(1.10), which was also established in $[5,15]$ by using completely different approaches.

Corollary 2.1. Let $Q^{0}, \tilde{Q} \in C^{2+\alpha}(\bar{\Omega})$. For any fixed $T>0$, suppose $\left\|Q^{0}\right\|_{L^{\infty}(\Omega)},\|\tilde{Q}\|_{L^{\infty}(\partial \Omega)}$ and $\frac{a^{-}}{c}$ are sufficiently small. Then there exists a unique solution $Q(x, t)$ to the problem (1.9)-(1.10), with the following properties:

$$
Q \in L^{\infty}\left(0, T ; L^{\infty}(\Omega) \cap H^{1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{2}(\Omega)\right), \text { and } Q(x, t) \in \mathcal{S}^{(2)}, \forall(x, t) \in \Omega \times[0, T]
$$

Further, $\|Q\|_{L^{\infty}(\Omega)}$ always stays small during evolution.
It is worth mentioning that the regularity in Corollary 2.1 can be improved using bootstrap argument so that the weak solution $Q$ is indeed a classical solution.

Next let us recall Lemma 3.2 in [15] that relates to the continuous dependence on the initial data.

Lemma 2.2. Let

$$
Q_{i} \in L^{\infty}\left(0, T ; L^{\infty}(\Omega) \cap H^{1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{2}(\Omega)\right) \quad(i=1,2)
$$

be two global weak solutions to the problem (1.9)-(1.10) on ( $0, T$ ), with initial data $Q_{1}, Q_{2} \in L^{\infty}(\Omega) \cap$ $H^{1}(\Omega)$. Suppose $\left\|Q_{i}\right\|_{L^{\infty}(\Omega)}(i=1,2)$ are sufficiently small. Then for any $t \in(0, T)$, we have

$$
\begin{equation*}
\left\|\left(Q_{1}-Q_{2}\right)(t)\right\|_{L^{2}(\Omega)} \leq C e^{C t}\left\|Q_{01}-Q_{02}\right\|_{L^{2}(\Omega)} \tag{2.35}
\end{equation*}
$$

where $C>0$ is a constant that depends on $\Omega, Q_{0 i}(i=1,2), \tilde{Q}$ and the coefficients of the system, but not $t$.

By virtue of Lemma 2.2, we may relax the regularity assumption on the initial data $Q^{0}$, and henceforth we state the existence result as follows

Corollary 2.2. Let $Q^{0} \in H^{1}(\Omega) \cap L^{\infty}(\Omega), \tilde{Q} \in C^{2+\alpha}(\bar{\Omega})$. For any fixed $T>0$, suppose $\left\|Q^{0}\right\|_{L^{\infty}(\Omega)}$, $\|\tilde{Q}\|_{L^{\infty}(\partial \Omega)}$ and $\frac{a^{-}}{c}$ are sufficiently small. Then there exists a unique global weak solution $Q(x, t)$ to the problem (1.9)-(1.10) that satisfies

$$
Q \in L^{\infty}\left(0, T ; L^{\infty}(\Omega) \cap H^{1}(\Omega)\right) \cap L^{2}\left(0, T ; H^{2}(\Omega)\right), \quad Q(x, t) \in \mathcal{S}^{(2)}, \forall(x, t) \in \Omega \times[0, T]
$$

Further, the smallness of the $L^{\infty}$ norm of $Q$ is preserved during evolution
Proof. For $Q^{0} \in L^{\infty}(\Omega) \cap H^{1}(\Omega)$, let us use the standard mollifier to establish $Q^{\varepsilon, 0} \in C^{2+\alpha}(\varepsilon \rightarrow 0)$ with $Q^{\varepsilon, 0} \rightarrow Q^{0}$ in $H^{1}(\Omega)$, and $\left\|Q^{\varepsilon, 0}\right\|_{L^{\infty}} \leq\left\|Q^{0}\right\|_{L^{\infty}}$. Then $Q^{\varepsilon}(t)$ is the corresponding solution with initial data $Q^{\varepsilon, 0}$. As $Q^{\varepsilon} \in L\left(0, T ; L^{\infty} \cap H^{1}\right) \cap L^{2}\left(0, T ; H^{2}\right)$ and such bounds depend on the $L^{\infty} \cap H^{1}$ bound of $Q^{0}$ only, $Q^{\varepsilon}$ is a Cauchy sequence in $L^{\infty}\left(0, T ; L^{2}(\Omega)\right)$. Hence we define $Q(x, t)=\lim _{\varepsilon \rightarrow 0} Q^{\varepsilon}(x, t)$ that solves the equation weakly. Then we may proceed as before and the proof is complete.

Remark 2.2. It is pointed out in Theorem 2.2 that the initial data $Q^{0}$ of the evolution problem (1.9)-(1.10) can be relaxed from $C^{2+\alpha}$ to $H^{1} \cap L^{\infty}$. Regarding the boundary data $\tilde{Q}$, however, it seems that we cannot relax its regularity because of the Schauder estimates used in Proposition 2.1. On the other hand, one may easily find that it suffices to assume $\tilde{Q} \in C^{0}(\partial \Omega)$ to perform the maximum principle argument in Lemma 2.1.

## 3 Numerical experiments

We have shown in the previous section that the proposed numerical scheme preserves the symmetric and traceless properties of the tensor $Q^{n}(n \geq 1)$, provided the initial state $Q^{0}$ and boundary value are in the Q -tensor space $\mathcal{S}^{(2)}$. By parameterizing $Q$ as

$$
Q(\cdot, t)=\left(\begin{array}{cc}
p(\cdot, t) & q(\cdot, t)  \tag{3.1}\\
p(\cdot, t) & -q(\cdot, t)
\end{array}\right), \quad Q(\cdot, 0)=Q^{0}=\left(\begin{array}{cc}
p^{0} & q^{0} \\
p^{0} & -q^{0}
\end{array}\right),
$$

the numerical scheme (2.1)-(2.2) can be rewritten as:

$$
\begin{aligned}
\frac{p^{n+1}-p^{n}}{\Delta t}= & \zeta \Delta p^{n+1}-a p^{n+1}-2 c\left(\left|p^{n+1}\right|^{2}+\left|q^{n+1}\right|^{2}\right) p^{n+1}-2 L\left(p^{n+1}-p^{n}\right)\left(\left|\nabla p^{n+1}\right|^{2}+\left|\nabla q^{n+1}\right|^{2}\right) \\
& +2 L_{4}\left(p^{n} \partial_{x x} p^{n+1}-p^{n} \partial_{y y} p^{n+1}+2 q^{n} \partial_{x y} p^{n+1}+\partial_{x} p^{n} \partial_{x} p^{n+1}-\partial_{y} p^{n} \partial_{y} p^{n+1}\right) \\
& +L_{4}\left(2 \partial_{x} q^{n} \partial_{y} p^{n+1}+2 \partial_{y} q^{n} \partial_{x} p^{n+1}+\left|\partial_{y} p^{n+1}\right|^{2}+\left|\partial_{y} q^{n+1}\right|^{2}-\left|\partial_{x} p^{n+1}\right|^{2}-\left|\partial_{x} q^{n+1}\right|^{2}\right), \\
\frac{q^{n+1}-q^{n}}{\Delta t}= & \zeta \Delta q^{n+1}-a q^{n+1}-2 c\left(\left|p^{n+1}\right|^{2}+\left|q^{n+1}\right|^{2}\right) q^{n+1}-2 L\left(q^{n+1}-q^{n}\right)\left(\left|\nabla p^{n+1}\right|^{2}+\left|\nabla q^{n+1}\right|^{2}\right) \\
& +2 L_{4}\left(p^{n} \partial_{x x} q^{n+1}-p^{n} \partial_{y y} q^{n+1}+2 q^{n} \partial_{x y} q^{n+1}+\partial_{x} p^{n} \partial_{x} q^{n+1}-\partial_{y} p^{n} \partial_{y} q^{n+1}\right) \\
& +2 L_{4}\left(\partial_{x} q^{n} \partial_{y} q^{n+1}+\partial_{y} q^{n} \partial_{x} q^{n+1}-\partial_{x} p^{n+1} \partial_{y} p^{n+1}-\partial_{x} q^{n+1} \partial_{y} q^{n+1}\right) .
\end{aligned}
$$

We now describe briefly our numerical approach. For simplicity of implementation, we consider the periodic boundary conditions and use the Fourier spectral method [12] for the space variable. Thus at each time step, we have a coupled nonlinear system for the Fourier approximation of


Figure 1: Temporal error $e_{2}(t=0.5)$ (left) and $e_{\infty}(t=0.5)$ (right) for Example 1.
$\left(p^{n+1}, q^{n+1}\right)$, which will be solved by using the Newton iteration method. At each Newton iteration, we need to solve a coupled linearized system. These linearized systems always have non-constant coefficients that make a direct solution by Fourier spectral method difficult and expensive. Therefore, we solve them by using the preconditioned BiCGSTAB method with a preconditioner coming from a suitable linear system with constant coefficients for which the Fourier spectral method reduces to a diagonal system. Hence, the cost of each BiCGSTAB iteration is simply a matrix-vector product which can be done in $O\left(N^{2} \log N\right)$ ( $N$ being the number of modes in each direction) operations with a pseudo-spectral matrix-free approach using FFT [12, 26].

We now present some numerical results obtained by using the above approach.
Example 1. (Accuracy test) We set $\Omega=[-2,2] \times[-2,2]$ and take the initial data to be

$$
\begin{equation*}
p^{0}(x)=\sin \left(\pi x_{1} / 2\right) \sin \left(\pi x_{2} / 2\right), \quad q^{0}(x)=\cos \left(\pi x_{1} / 2\right) \cos \left(\pi x_{2} / 2\right), \quad x=\left(x_{1}, x_{2}\right)^{T} \in \Omega . \tag{3.2}
\end{equation*}
$$

The other parameters are given as

$$
\begin{equation*}
\zeta=2, \quad a=0.5, \quad c=4, \quad L_{4}=0.1 . \tag{3.3}
\end{equation*}
$$

Since we do not know the explicit form of the exact solution, we take the 'reference' solution $\left(p\left(\cdot, t_{n}\right), q\left(\cdot, t_{n}\right)\right)$ to be the numerical solution obtained by using the proposed scheme with the stabilizing constant $L=0.5$, space mesh size $h_{e}=1 / 32$ which well resolves the solution, and a small time step $\tau_{e}=10^{-4}$.

We first look at the the temporal errors. We take the space mesh size $h=1 / 32$ such that the spatial errors are negligible. Let $\left(p_{\tau}^{n}, q_{\tau}^{n}\right)$ be the numerical approximations obtained by our scheme


Figure 2: (Example 2) Orientation of liquid crystal at different time $t$.
at $t=t_{n}$ with $h=1 / 32$ and time step $\tau$, and we introduce the $L^{2}$ and $L^{\infty}$ error functions as

$$
e_{2}\left(t_{n}\right)=\sqrt{\left\|p_{\tau}^{n}-p_{\tau_{e}}^{n}\right\|_{L^{2}}^{2}+\left\|q_{\tau}^{n}-q_{\tau_{e}}^{n}\right\|_{L^{2}}^{2}}, \quad e_{\infty}\left(t_{n}\right)=\left\|\sqrt{\left|p_{\tau}^{n}-p_{\tau_{e}}^{n}\right|^{2}+\left|q_{\tau}^{n}-q_{\tau_{e}}^{n}\right|^{2}}\right\|_{\infty}
$$

Fig. 1 shows the temporal errors for different stabilizing constant $L$. It is clear that the scheme is first order accurate in time.

Example 2. We choose $\Omega=[-2,2] \times[-2,2]$ with periodic boundary conditions and $\zeta=0.4, \quad a=$ $-2, \quad c=1, \quad L_{4}=0.08, \quad L=0.1$. We set the initial state

$$
\begin{equation*}
Q^{0}(x)=s^{0}(x)\left(2 \vec{n}^{0} \otimes \vec{n}^{0}-\mathbf{I}_{2}\right) \tag{3.4}
\end{equation*}
$$



Figure 3: (Example 3) Orientation of liquid crystal at different time $t$.
with $s^{0}(x)=0.1$ and

$$
\vec{n}^{0}(x)= \begin{cases}(1,0)^{t}, & x \in[-1,1] \times[-1,1] ;  \tag{3.5}\\ (0,1)^{t}, & \text { otherwise }\end{cases}
$$

being the unit vector in $\mathbb{R}^{2}$ representing the direction of the liquid crystal at position $x$.
Example 3. We choose the same parameters as in Example 2 but with the initial state

$$
\begin{equation*}
Q^{0}(x)=s^{0}(x)\left(2 \vec{n}^{0} \otimes \vec{n}^{0}-\mathbf{I}_{2}\right), \tag{3.6}
\end{equation*}
$$



Figure 4: Evolutions of $|Q|^{2}$ and the energy for Examples 2\& 3.
with $s^{0}(x)=0.1$ and

$$
\vec{n}^{0}(x)= \begin{cases}(1,0)^{t}, & x \in[-1.5,1.5] \times[-1.5,1.5]  \tag{3.7}\\ (0,1)^{t}, & \text { otherwise }\end{cases}
$$

In the computations for Examples 2 and 3, we choose $\tau=0.0025$ and $h=1 / 32$. Figs. 2 and 3 show the orientation of the liquid crystal during the time evolution. We observe from Fig. 2 that the final steady states depend on the initial data. For Example 2 (cf. Fig. 2), initially there are more vertical molecules than horizontal molecules. The set with horizon molecules shrinks with time, and the liquid crystal directions eventually approach to the uniform vertical configuration. On the other hand, for Example 3 (cf. Fig. 3), there are more horizontal molecules than vertical molecules at $t=0$. The set of horizontal molecules expands towards boundary while its shape oscillates, and eventually and the liquid crystal directions eventually approach to the uniform horizontal configuration.

Next, we examine the time evolution of $L^{\infty}$ norm of $|Q|^{2}$ and bulk energy $\mathcal{F}_{\text {bulk }}(1.4)$, see Fig. 4. We observe that, when the $L^{\infty}$ bound of the Q-tensor order parameter is sufficiently small, the elliptic part in the equation will force the system approach to a uniform state, and $|Q|^{2}$ will approach to the minimizer of the bulk energy $\mathcal{F}_{\text {bulk }}(1.4)$, which is constant 2 . In all our numerical results, the $L^{\infty}$ norm of the numerical solutions remains to be small for small initial data, as proved in the analysis.

Finally we examine the computational effectiveness of our approach by looking at the convergences of Newton iteration and total BiCGSTAB iterations at each time step during the evolution of Example 2. Fig. 5 displays the number of Newton iterations and total BiCGSTAB iterations per time step, with tolerance $10^{-12}$ for the Newton iterations and $10^{-10}$ for the BiCGSTAB itera-


Figure 5: Numbers of Newton iteration and BiCGSTAB iterations at each time step for Example 2 with stabilizing constant $L=0.1$ (left) and $L=10$ (right). The tolerance of absolute error for Newton iteration is $10^{-12}$ (measured in maximum norm of the residue) and the tolerance of relative error for BiCGSTAB is $10^{-10}$ (measured in $l^{2}$ norm of the residue).
tions. We observe that the number of Newton iterations per time step ranges between 1-5, and the total BiCGSTAB iterations per time step ranges between 2-11, which indicates that, on average, the BiCGSTAB converges in just 2-3 iterations for each linearized system. These results indicate that our numerical approach is very efficient.

## 4 Conclusion

In this paper, we proposed an unconditionally stable numerical scheme to solve a 2 D Q-tensor model for liquid crystal, and established its unique solvability and convergence. As a byproduct of our analysis, we also established the well-posedness of the original PDE system for the 2D Q-tensor model, which has been shown previously with completely different approaches.

The main difficulty in the analysis came from an unusual cubic $L_{4}$ term in the elastic energy, which made the free energy unbounded from below and caused great challenges in both analysis and computation. By adding a stabilized term in our scheme, we were able to show that the $L^{\infty}$ norm of the numerical solution can be kept small which guaranteed the stability and the well-posedness. Numerical tests showed that the scheme is indeed first order accurate for a wide range of stabilizing constants, and produces physically consistent numerical simulations.

We only discussed a 2D Q-tensor model in this paper. Extensions to the 3D case, as well as the full dynamical model coupled with Navier-Stokes equations entail significant analytical difficulties,
and they will be considered in our future work.

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